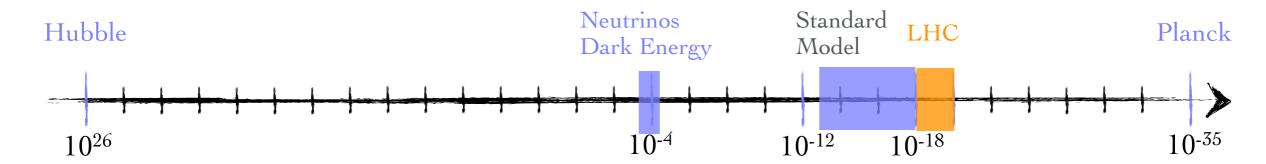
# The case for Ultra-Light Boson Dark Matter

Asimina Arvanitaki Perimeter Institute

# The High Energy Frontier



#### Particle Physics — Precision vs Energy Frontier



Scale in meters

#### Opportunities to probe the low energy frontier

Atom Interferometry

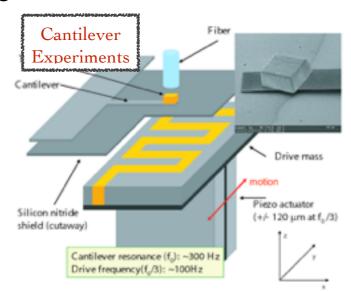
Atomic

Clocks

- Tests of Gravity
- Gravitational Wave detection at low frequencies
- Tests of Atom Neutrality at 30 decimals
- Dimopoulos, Geraci (2003) Dimopoulos, Kasevich et. al.(2006-2008)

- •Setting the Time Standard
- Dilaton Dark MatterDetection
- AA, Huang, Van Tilburg (2014)

- Short Distance Tests of Gravity
- •Extra Dimensions



Dimopoulos, Kapitulnik (1997)

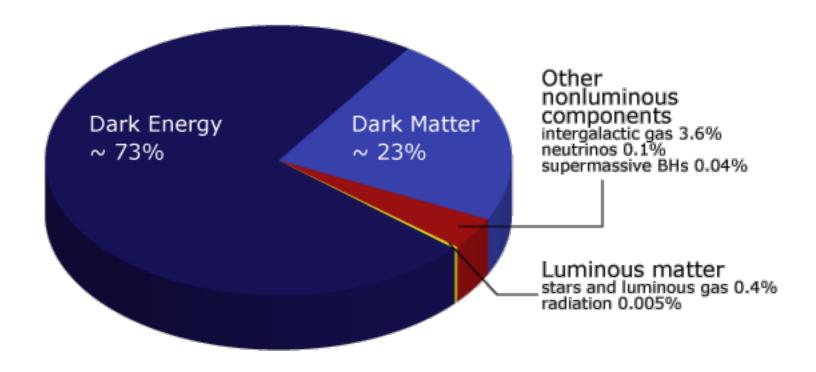
- Axion Dark MatterDetection
- •Axion Force

Detection



Graham et. al. (2012) AA, Geraci (2014)

## The Mystery of Dark Matter



#### Models of Dark Matter

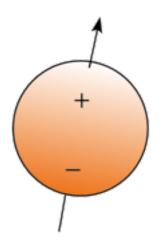
• What is it made out of?

• How is it produced?

• Does it have interactions other than gravitational?

## Why is the Electric Dipole Moment of the Neutron Small?

The Strong CP Problem and the QCD axion



Neutron Electric Dipole Moment  $\sim$  e fm  $\theta_{QCD}$ 

$$L_{\scriptscriptstyle ext{SM}} \supset rac{g_s^2}{32\pi^2} heta_{\scriptscriptstyle ext{QCD}} G^a ilde{G}^a$$

Experimental bound:  $\theta_{QCD}$  <  $10^{-10}$ 

 $Solution: \\ \theta_{QCD} \ is \ a \ dynamical \ field, \ an \ axion$ 

Weinberg (1978) and Wilczek (1978) Peccei and Quinn (1977)

Axion mass from QCD:

$$\mu_a \sim 6 \times 10^{-11} \text{ eV} \frac{10^{17} \text{ GeV}}{f_a} \sim (3 \text{ km})^{-1} \frac{10^{17} \text{ GeV}}{f_a}$$

fa: axion decay constant

Extra dimensions

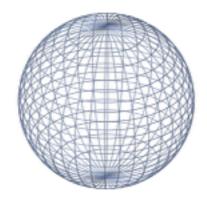
• Extra dimensions

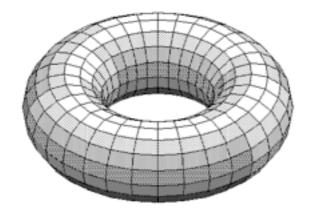
• Gauge fields

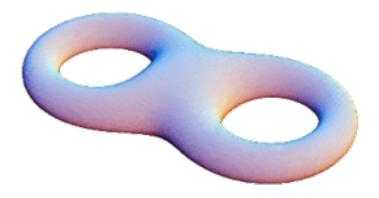
• Extra dimensions

• Gauge fields

● Topology



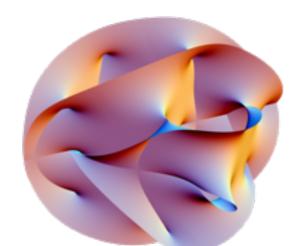




• Extra dimensions

• Gauge fields

● Topology



• Extra dimensions

• Gauge fields

● Topology



Give rise to a plenitude of Universes

Extra dimensions

• Gauge fields

● Topology



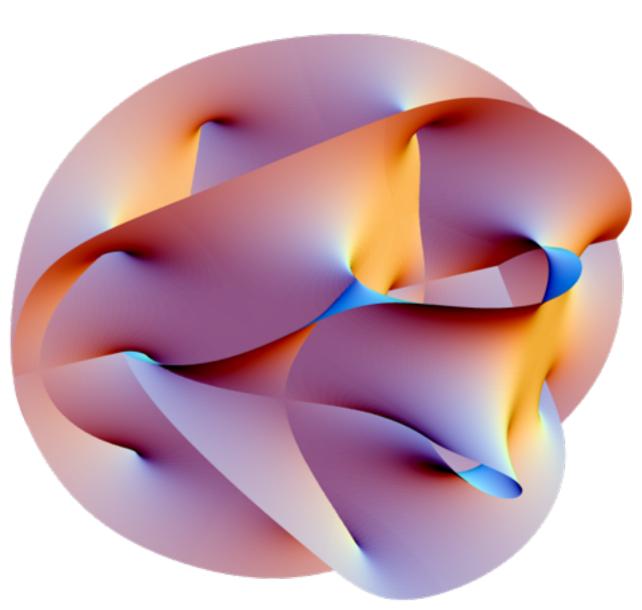
Give rise to a plenitude of massless particles in our Universe

#### A Plenitude of Massless Particles

Compactification naturally gives rise to massless particles

In the presence of non-trivial topology,
non-trivial gauge field configurations can carry no
energy,

resulting in 4D massless particles



## Non-trivial gauge configurations

#### The Aharonov-Bohm Effect





Taking an electron around the solenoid

$$e \int A_{\mu} dx^{\mu} = e \times \text{Magnetic Flux}$$

while

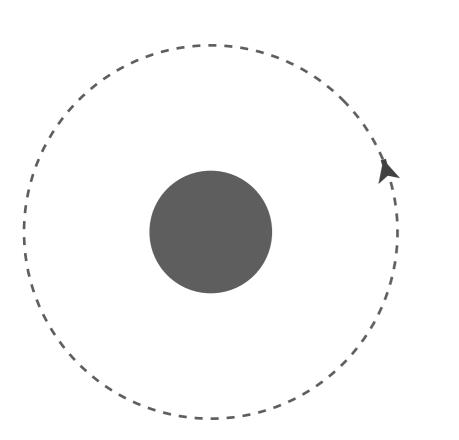
$$\vec{B} = 0$$

Energy stored only inside the solenoid

Non-trivial gauge configuration far away carries no energy

## Non-trivial gauge configurations

#### The Aharonov-Bohm Effect



Taking an electron around the solenoid

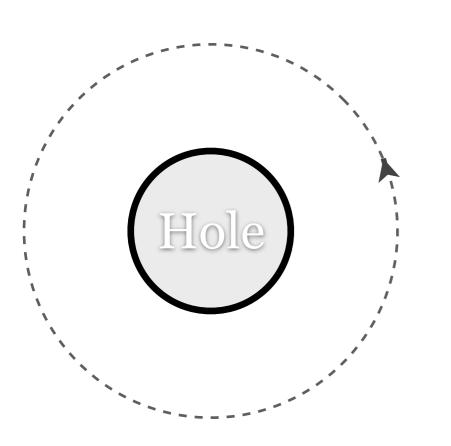
$$e\int A_{\mu}dx^{\mu}=e imes {
m Magnetic}$$
 Flux 
$${
m while}$$
 
$$\vec{B}=0$$

Energy stored only inside the solenoid

Non-trivial gauge configuration far away carries no energy

#### Non-trivial gauge configurations

#### The Aharonov-Bohm Effect



Taking an electron around the solenoid

$$e \int A_{\mu} dx^{\mu} = e \times \text{Magnetic Flux}$$
 while

$$\vec{B} = 0$$

Non-trivial topology:

"Blocking out" the core still leaves a non-trivial gauge, but no mass

#### A Plenitude of Massless Particles

• Spin-0 non-trivial gauge field configurations: String Axiverse

• Spin-1 non-trivial gauge field configurations: String Photiverse

• Fields that determine the shape and size of extra dimensions as well as values of fundamental constants: Dilatons, Moduli, Radion

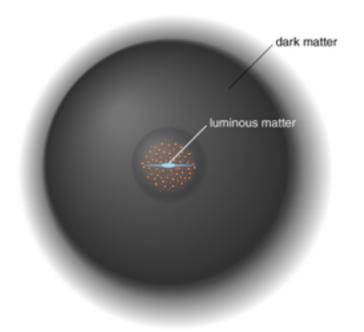
# Properties of Plenitude of Particles from String Theory

• They couple very weakly to the Standard Model

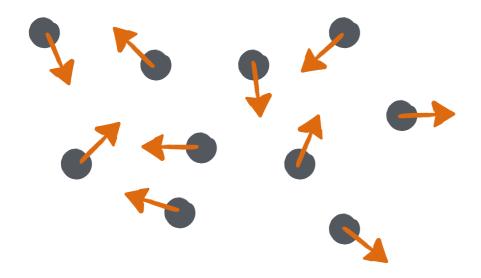
• They can be extremely light

 Constrained if the coupling is large enough by astrophysics, BBN, CMB...

Dark Matter Particles in the Galaxy

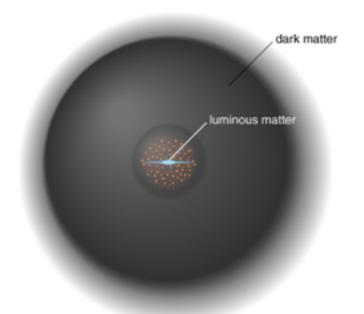


Usually we think of ...

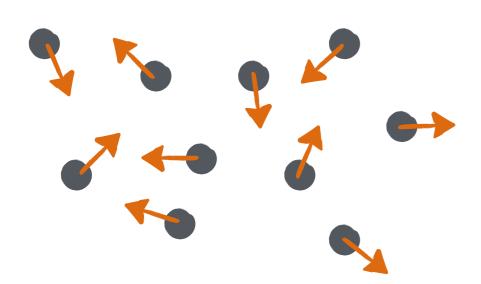


like a WIMP

Dark Matter Particles in the Galaxy

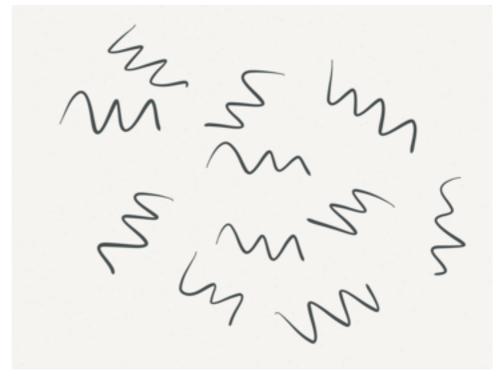


Usually we think of ...



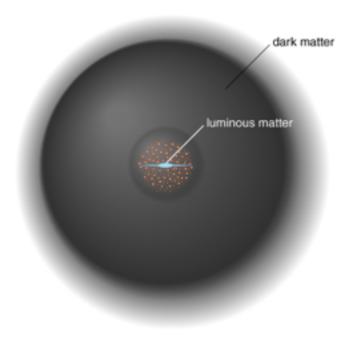
like a WIMP

instead of...



$$\lambda_{DM} = \frac{\hbar}{m_{DM}v}$$

Dark Matter Particles in the Galaxy



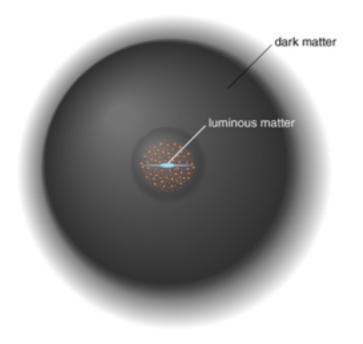
Decreasing DM Mass



$$\lambda_{DM} = \frac{\hbar}{m_{DM}v}$$



Dark Matter Particles in the Galaxy



Decreasing DM Mass



$$\lambda_{DM} = \frac{\hbar}{m_{DM}v}$$



Equivalent to a Scalar wave

#### Going from DM particles to a DM "wave"



When 
$$n_{DM} > \frac{1}{\lambda_{DM}^3}$$

In our galaxy this happens when  $m_{DM} < 1 \text{ eV/c}^2$ 

we can talk about DM  $\phi(x,t)$  and locally

$$\phi(t) \approx \phi_0 \cos \omega_{DM} t$$

with amplitude

$$\phi_0 \propto \frac{\sqrt{\mathrm{DM density}}}{\mathrm{DM mass}}$$

with frequency

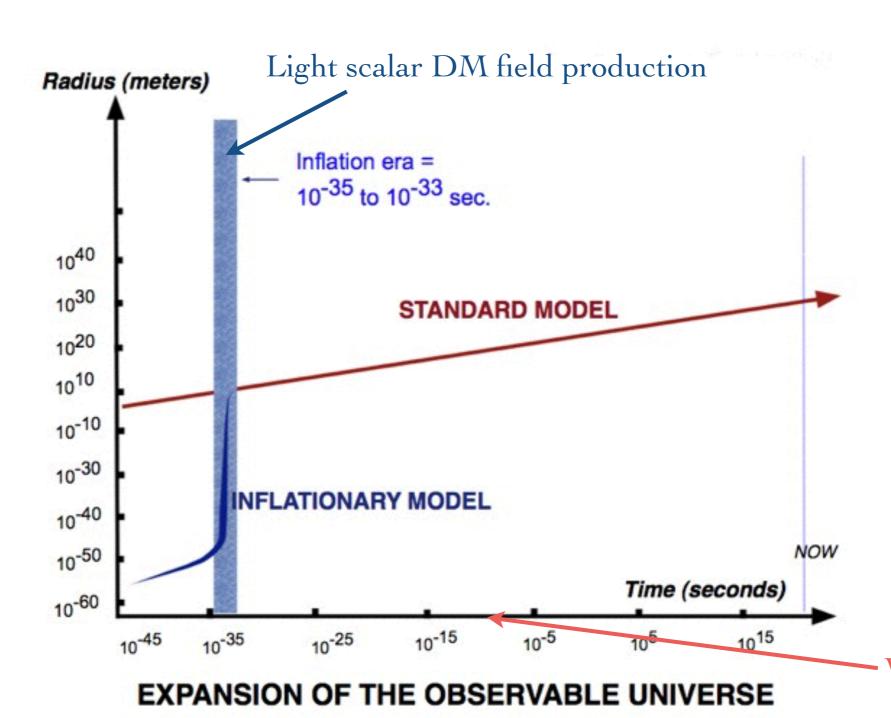
$$\omega_{DM} pprox \frac{m_{DM}c^2}{\hbar}$$

and finite coherence

$$\delta\omega_{DM} \approx \frac{m_{DM}v^2}{\hbar} = 10^{-6}\omega_{DM}$$

#### Scalar DM field Production Mechanism

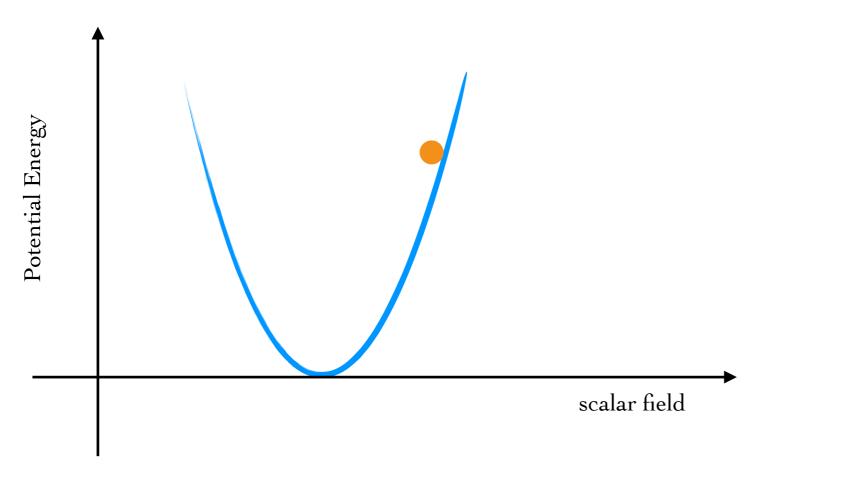
• The "misalignment mechanism" during inflation



WIMP production

#### Light Scalar Dark Matter

• Produced by the misalignment mechanism



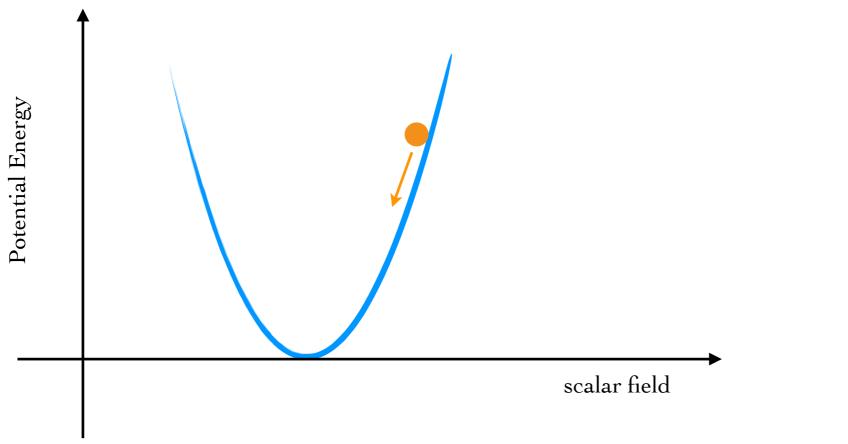
Frozen when: Hubble >  $m_{\phi}$ 

Initial conditions set by inflation

\*The story changes slightly if DM is a dark photon

#### Light Scalar Dark Matter

Produced by the misalignment mechanism



Frozen when: Hubble >  $m_{\phi}$ 

Oscillates when: Hubble  $< m_{\phi}$ 

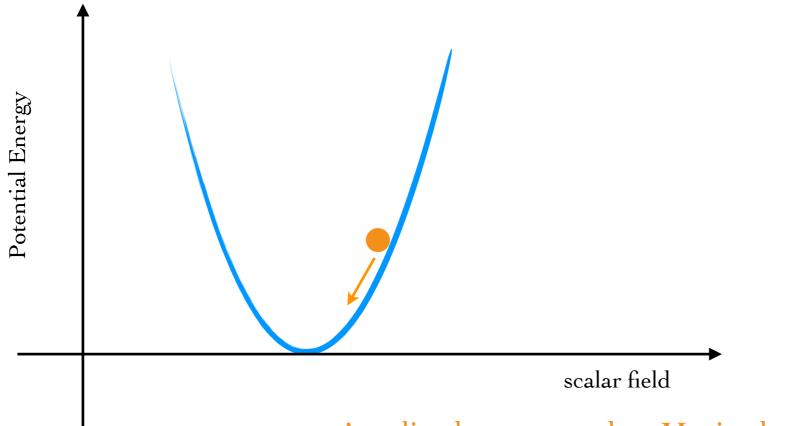
 $\rho_{\phi}$  scales as a<sup>-3</sup> just like Dark Matter

Initial conditions set by inflation

\*The story changes slightly if DM is a dark photon

#### Light Scalar Dark Matter Today

• If  $m_{\phi} < 1$  eV, can still be thought of as a scalar field today



$$m_{\phi}^2 \phi_o^2 \cos^2(m_{\phi} t) \sim \rho_{\phi}$$

Coherent for  $v_{vir}^{-2} \sim 10^6$  periods

Amplitude compared to M<sub>Pl</sub> in the galaxy:

$$\kappa \phi_0 = \frac{\sqrt{8\pi\rho_\phi}}{m_\phi M_{\rm Pl}} = 6.4 \cdot 10^{-13} \left(\frac{10^{-18} \,\text{eV}}{m_\phi}\right)$$

#### Axion Dark Matter

Axion wind (ex. CASPER, QUAX)

$$\mathcal{L} \supset rac{ec{
abla}\phi}{f_{\phi}} \cdot ec{\sigma}$$
 just like detecting a  $ec{B}_{eff} \equiv rac{ec{
abla}\phi}{\mu_f f_{\phi}}$ 

Axion-to-photon conversion (ex. ADMX)

$$B_{\phi} \sim g_{\phi\gamma\gamma}\phi B_{ext}$$

#### Dark Photon Dark Matter

• Abundance from inflation depends on  $H_I$ 

P. W. Graham et al (2015)

$$\Omega_A = \Omega_{
m cdm} imes \sqrt{rac{m}{6 imes 10^{-6}\,{
m eV}}} \left(rac{H_I}{10^{14}\,{
m GeV}}
ight)^2$$

Detected if kinetically mixed with the photon

$$\mathcal{L} \supset \epsilon F_{EM} F_{DM}$$

• Detected like a photon (ex. DM Radio) DM electric field = DM magnetic field  $\sim \sqrt{\rho_{DM}}$ 

#### Moduli Dark Matter

Moduli set values of measured fundamental constants

• Couple non-derivatively to the Standard Model (as well axions with CP violation)

Examples of couplings

$$\mathcal{L} = \mathcal{L}_{SM} + \sqrt{\hbar c} \frac{\phi}{\Lambda} \mathcal{O}_{SM}$$

$$\mathcal{O}_{SM} \equiv m_e e \bar{e}, \ m_q q \bar{q}, \ G_s^2, \ F_{EM}^2, \dots$$

#### Fundamental constants are not really constants

#### Oscillating Fundamental Constants

AA, J. Juang, K. Van Tilburg (2014)

From the local oscillation of Dark Matter

Ex. for the electron mass:

$$d_{m_e}\sqrt{\hbar c}\frac{\phi}{M_{Pl}}m_ec^2e\bar{e}$$

 $M_{\rm pl}$  =  $10^{18}~{
m GeV}$ reduced Planck scale in energy

$$\frac{\delta m_e}{m_e} \approx \frac{d_{m_e} \phi_0}{M_{Pl}} \cos(\omega_{DM} t)$$

$$= 6.4 \times 10^{-13} \cos(\omega_{DM} t) \left(\frac{10^{-18} \text{ eV}}{m_{DM} c^2}\right) \left(\frac{d_{m_e}}{1}\right)$$

d<sub>me</sub>: coupling strength relative to gravity

Fractional variation set by square root of DM abundance

Need an extremely sensitive probe...

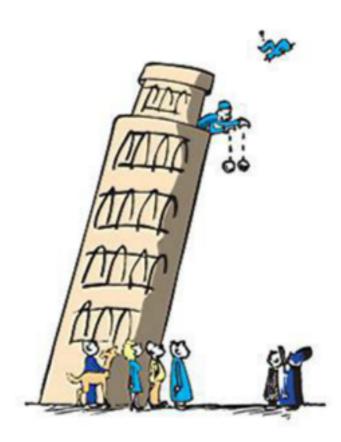
### Ultra-light Scalar Dark Matter

• Mediates new interactions in matter

• Generates a fifth force in matter



• Generates Equivalence Principle violation



#### Light Scalar Dark Matter Detection

Atomic Clocks searches

• Resonant-Mass Detector searches

Axion Force Experiments

• Detecting ultra-light particles with Astrophysical Black Holes

# Oscillating Atomic and Nuclear Energy Splittings due to Dark Matter

Optical Splittings

$$\Delta E_{
m optical} \propto lpha_{EM}^2 m_e \sim {
m eV}$$

Hyperfine Splittings

$$\Delta E_{
m hyperfine} \propto \Delta E_{
m optical} lpha_{EM}^2 \left(rac{m_e}{m_p}
ight) \sim 10^{-6} \, {
m eV}$$

Nuclear Splittings

$$\Delta E$$
 (m<sub>p</sub>,  $\alpha_s$ ,  $\alpha_{EM}$ )~ 1 MeV

#### Atomic Clocks

Kept tuned to an atomic energy level splitting

#### Current definition of a second:

the duration of 9192631770 periods of the radiation corresponding to the transition between the two hyperfine levels of the ground state of the caesium 133 atom

• Have shown stability of 1 part in 10<sup>18</sup>

Compared to 1 part in 10<sup>13</sup> expected by DM

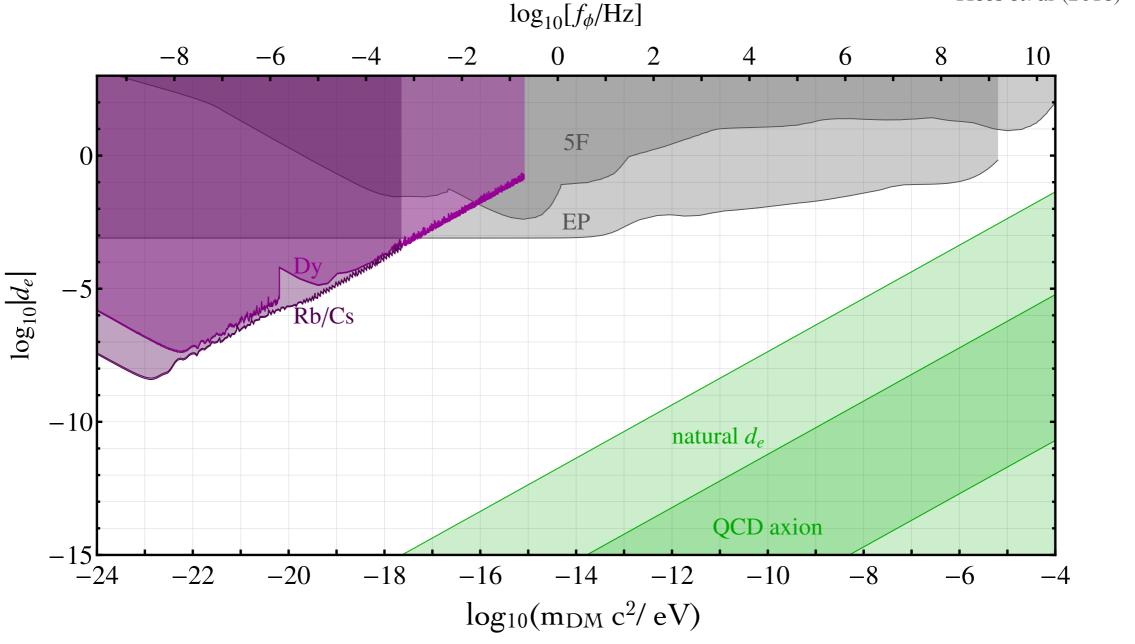
Have won several Nobel prizes in the past 20 years

## The Dy isotope and Rb/Cs Clock Comparison

sensitivity to  $\alpha_{EM}$  variations

Ken Van Tilburg and the Budker group (2015)

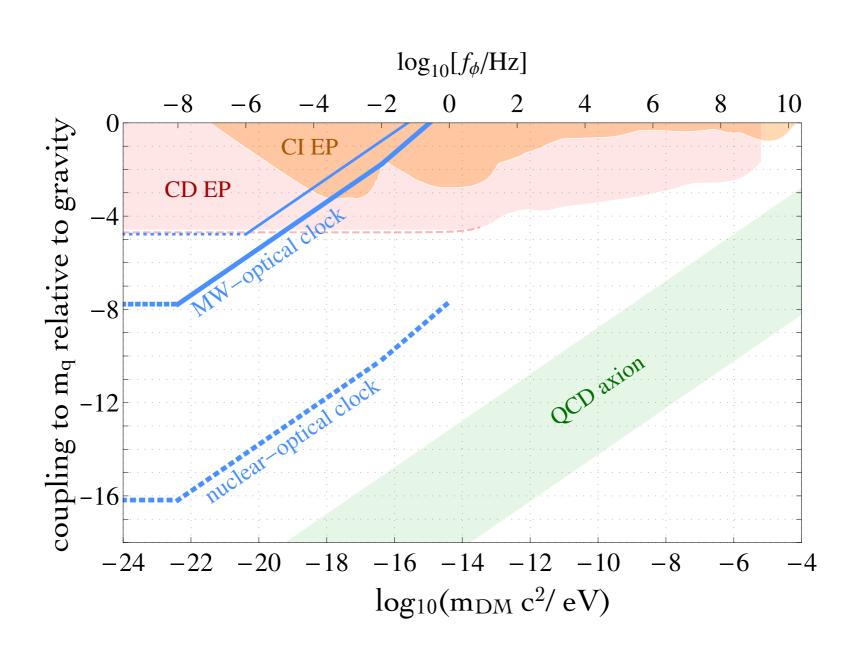
Hees et. al (2016)



Analysis performed with existing data

## Nuclear to Optical Clock Comparison

Future Sensitivity of a <sup>229</sup>Th clock



## Oscillating interatomic distances

• The Bohr radius changes with DM

• 
$$r_B \sim (\alpha m_e)^{-1}$$

$$\frac{\delta r_B}{r_B} = -\left(\frac{\delta \alpha_{EM}}{\alpha_{EM}} + \frac{\delta m_e}{m_e}\right)$$

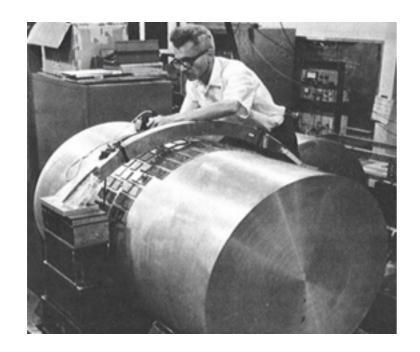
The size of solids changes with DM

• 
$$L \sim N (\alpha m_e)^{-1}$$
 
$$\frac{\delta L}{L} = -\left(\frac{\delta \alpha_{EM}}{\alpha_{EM}} + \frac{\delta m_e}{m_e}\right)$$

For a single atom  $\delta r_B \sim 10^{-30}$  m Need macroscopic objects to get a detectable signal

#### Resonant-Mass Detectors

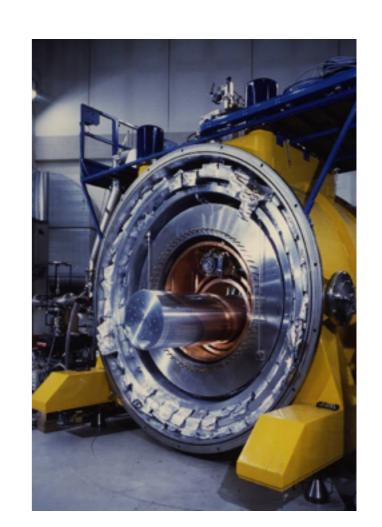
• In the 1960's: The Weber Bar



Fractional length variation  $\delta L/L\sim 10^{-17}$ 

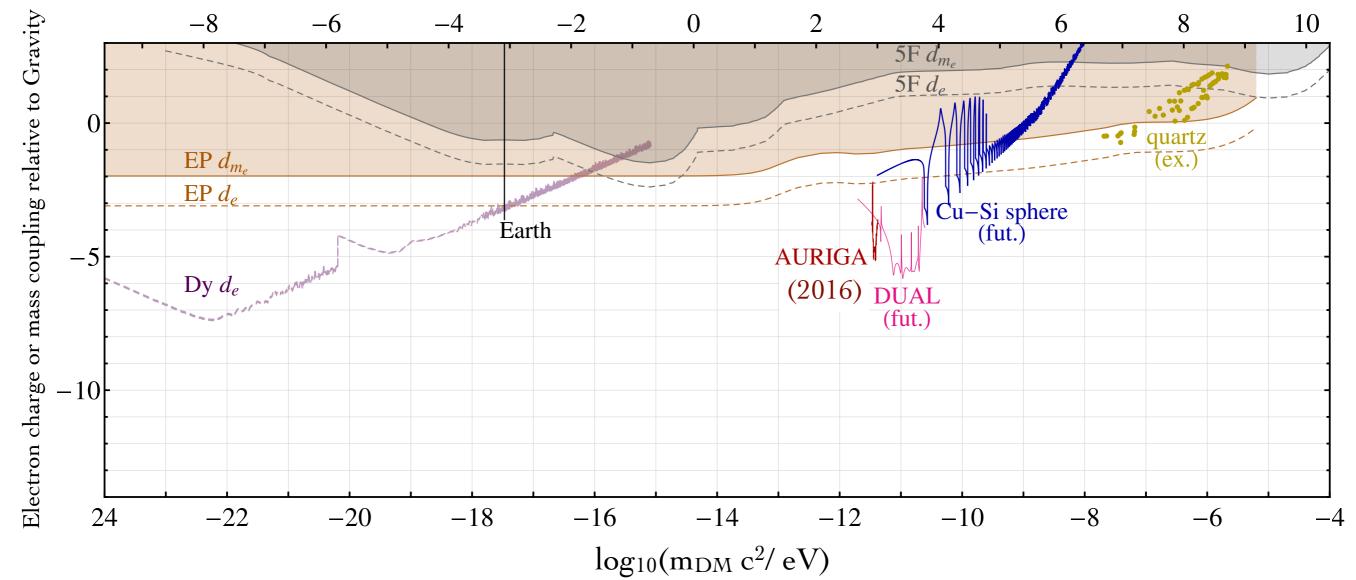
Today: AURIGA, NAUTILUS, MiniGrail

Fractional length variation  $\delta L/L{\sim}10^{\text{-}23}$ 



### What can be done in the future?

 $\log_{10}[f_{\phi}/\mathrm{Hz}]$ 

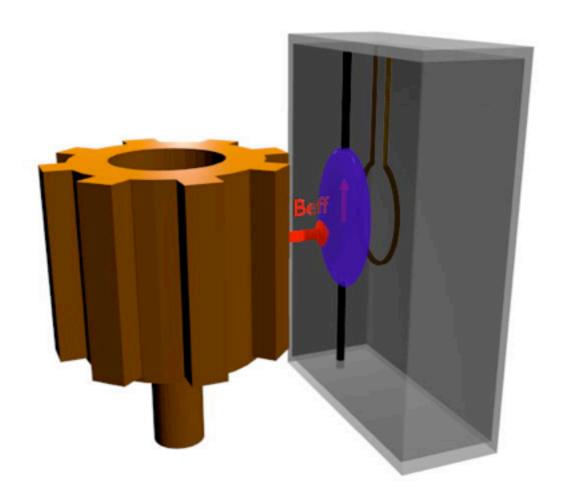


### Axion Resonant InterAction DetectioN Experiment

with Andrew Geraci (2014)

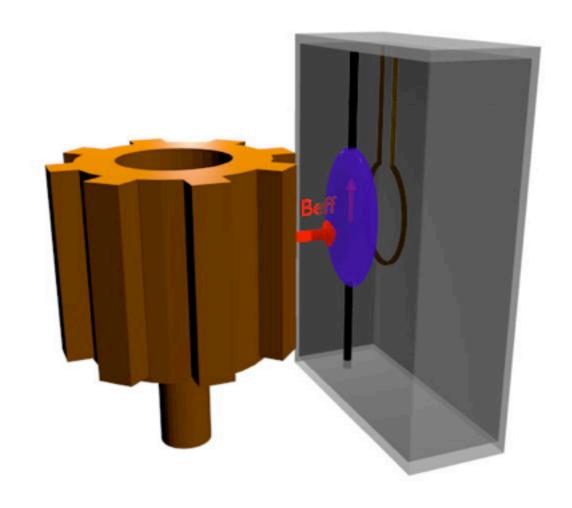
and A. Kapitulnik, Chen-Yu Liu, J. Long, Y. Semertzidis, M. Snow (to be built)

### Axion Resonant InterAction DetectioN Experiment



He-3 NMR sample with T<sub>2</sub> up to ~1000 sec

### Axion Resonant InterAction DetectioN Experiment

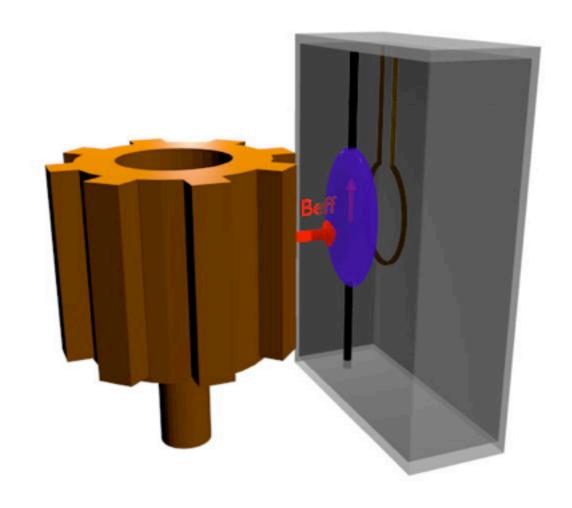


He-3 NMR sample with T<sub>2</sub> up to ~1000 sec

$$\begin{split} B_{\rm min} \approx p^{-1} \sqrt{\frac{2\hbar b}{n_s \mu_{^3{\rm He}} \gamma V T_2}} &= 3 \times 10^{-19}~{\rm T} \times \\ \left(\frac{1}{p}\right) \sqrt{\left(\frac{b}{1~{\rm Hz}}\right) \left(\frac{1~{\rm mm}^3}{V}\right) \left(\frac{10^{21}~{\rm cm}^{-3}}{n_s}\right) \left(\frac{1000~{\rm s}}{T_2}\right)} \end{split}$$

 $B_{min} = 10^{\text{-}16} \ T/(Hz)^{1/2}$  for SQUIDs

### Axion Resonant InterAction DetectioN Experiment



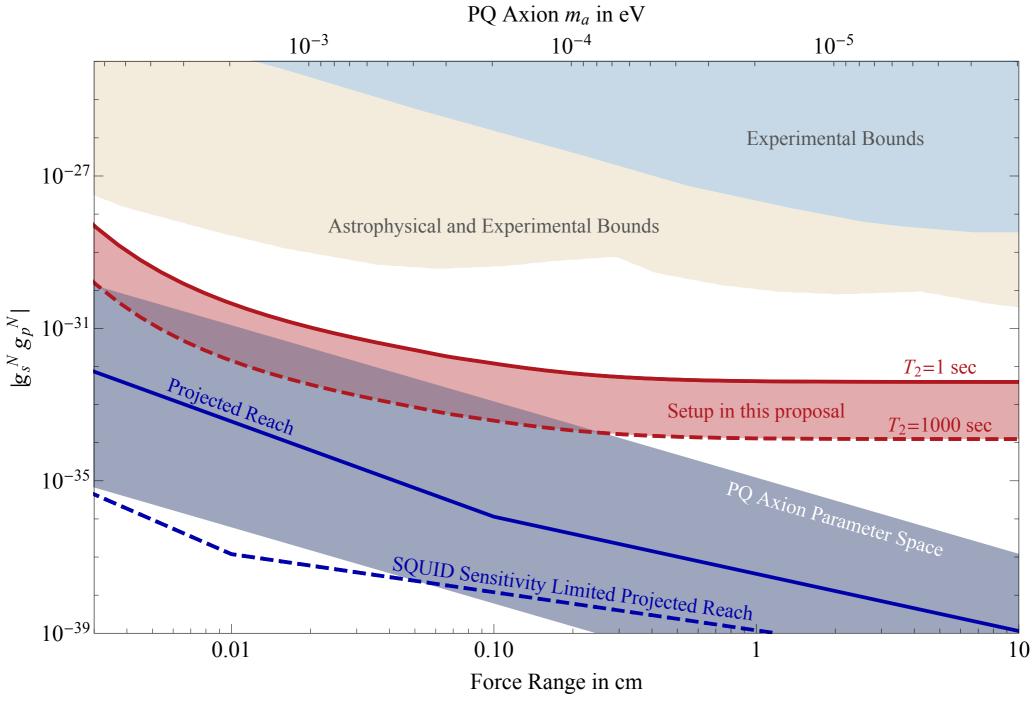
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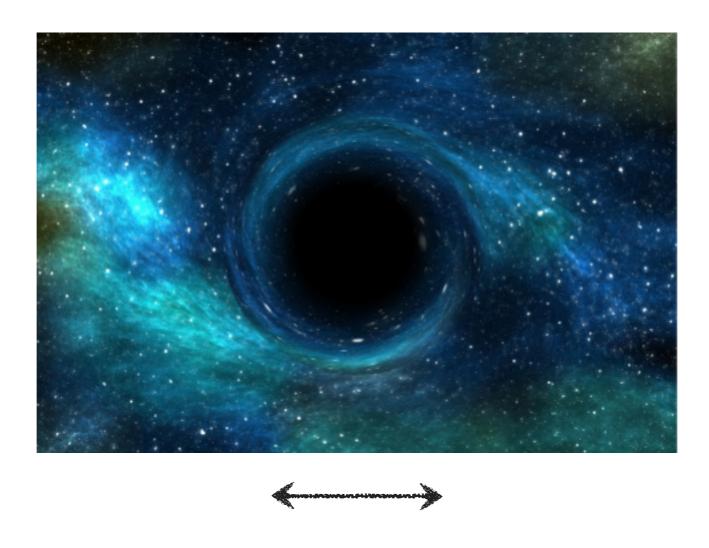
## Monopole-Dipole Interaction Reach

Unpolarized Source Mass with 10<sup>6</sup> sec integration



Projected Reach with increase of polarized spin density and larger NMR sample volume

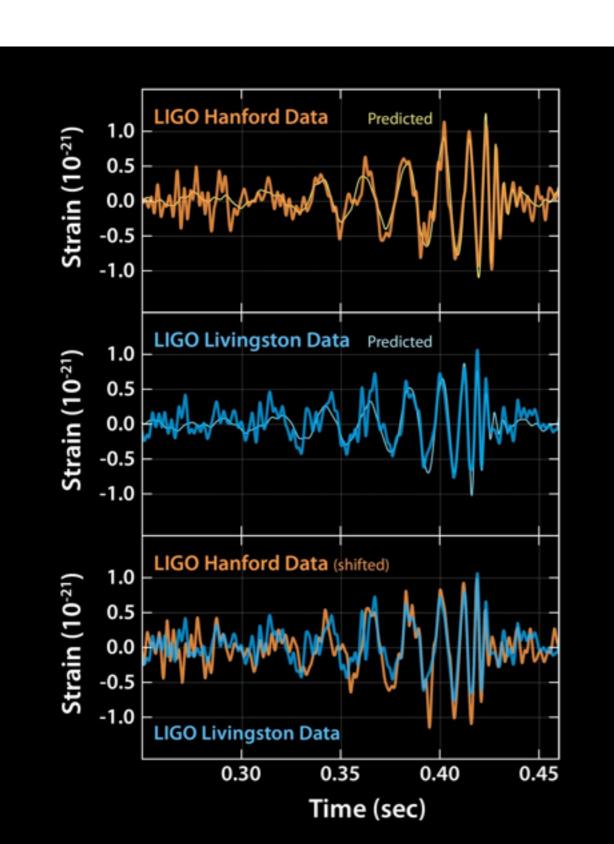
### Black Holes as Nature's Detectors

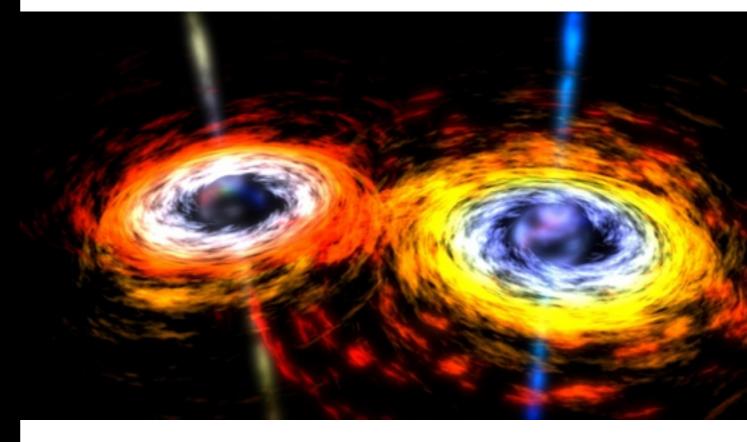


1 km -10 billion km

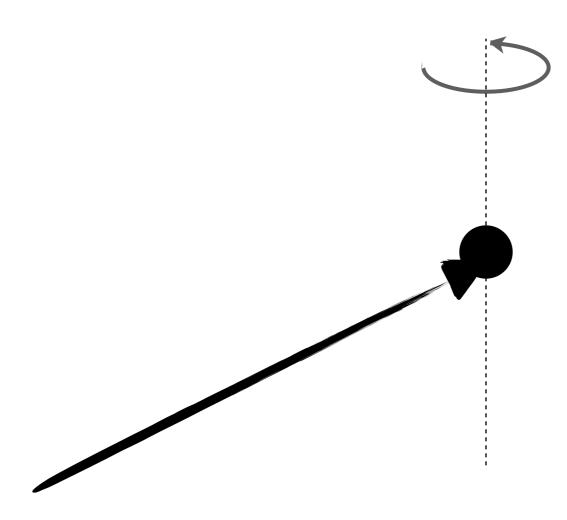
They can detect bosons of similar in size

## September 14, 2015



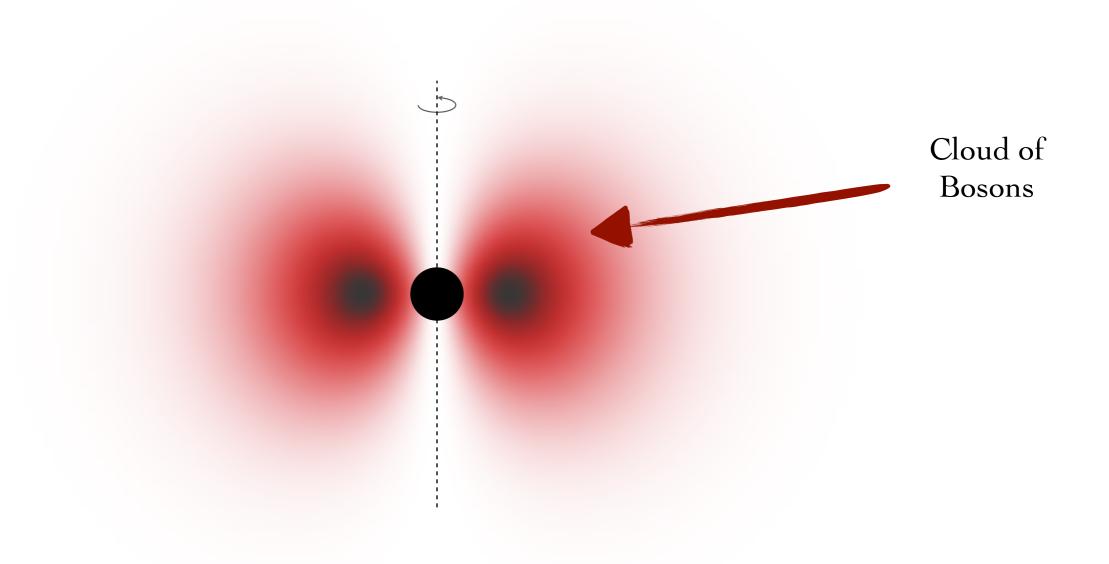


# Superradiance and The Gravitational Atom in the Sky



Rotating Black Hole

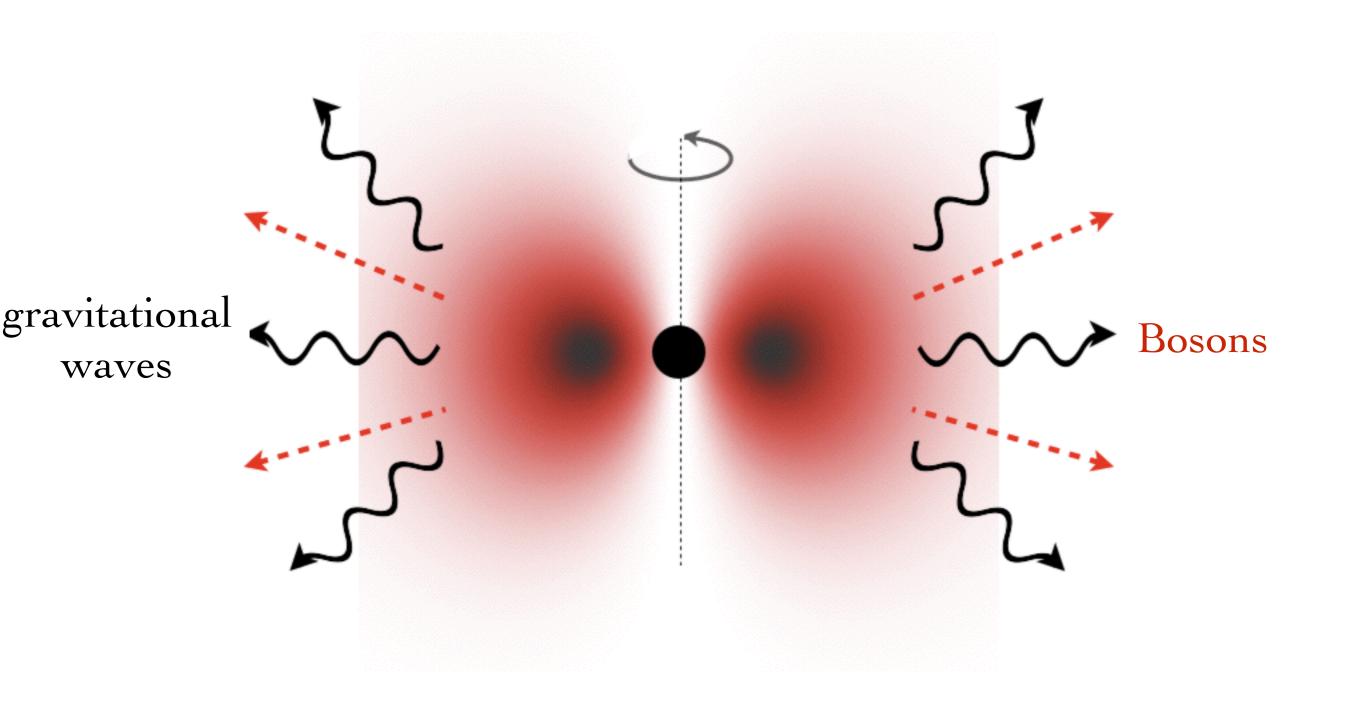
# Superradiance and The Gravitational Atom in the Sky



Particles Occupying the Same Bound Orbit:

Can be comparable to the number of protons in all the stars in the universe

# A Gravitational Wave Laser in The Sky



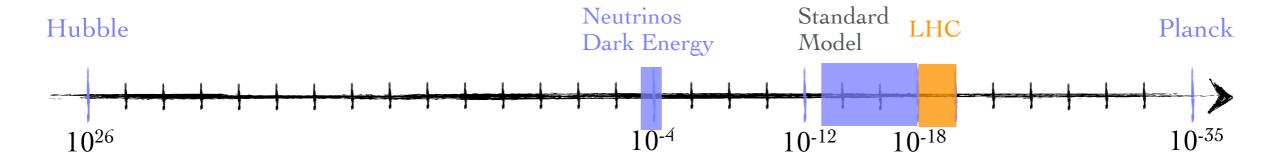
### Summary

• There is excellent theoretical motivation for boson DM candidates below 1 eV in mass

Correlated with observations across many experiments

This is only scratching the surface...

### The Precision Frontier



Scale in meters

There are more things in heaven and earth, Horatio, Than are dreamt of in your philosophy.

- Hamlet